Recent investigations on radiative shocks interacting with an obstacle

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Radiative shocks (RS) are phenomena widely observed in astrophysics, for example in supernovae remnants or supersonic jets. Moreover, massive stars (O and B stars) surrounding molecular clouds involve many radiation hydrodynamics processes. Here, we present experimental results obtained on highly radiative shocks (with a velocity up to 150 km/s) generated in a low-density gas filled cell obtained on the GEKKO XII laser facility. The RS was generated by using an ablator-pusher target (CH/Au/Ti), designed to limit as much as possible the preheating produced by the hot corona. The propagation media was Xe or He gas, with the aim to compare radiative effects in each medium.

Thanks to self-emission and visible probe diagnostics, we were able to analyze several aspects of the RS. First, a high electron density in the upstream region due to a radiative precursor is highlighted. Interaction between the radiative precursor and a solid obstacle (a quartz micro-balloon or an aluminum foil, mimicking a molecular cloud) has been studied both experimentally and theoretically. In addition, these results allow the characterization of the RS propagation, especially radiative effects and how they affect hydrodynamics.

References

[1] M. Koenig et al., Physics of Plasmas 24, 082707 (2017)[2] Th. Michel et al, HPLSE (accepted)